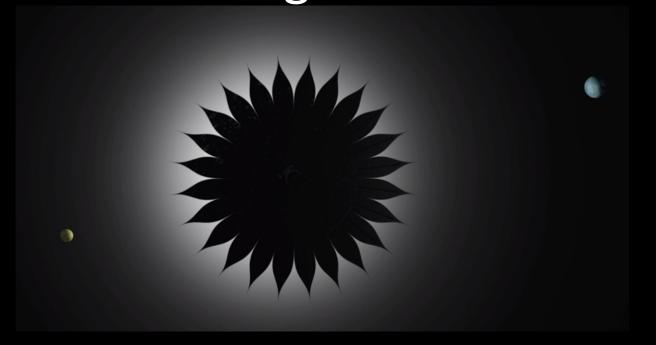
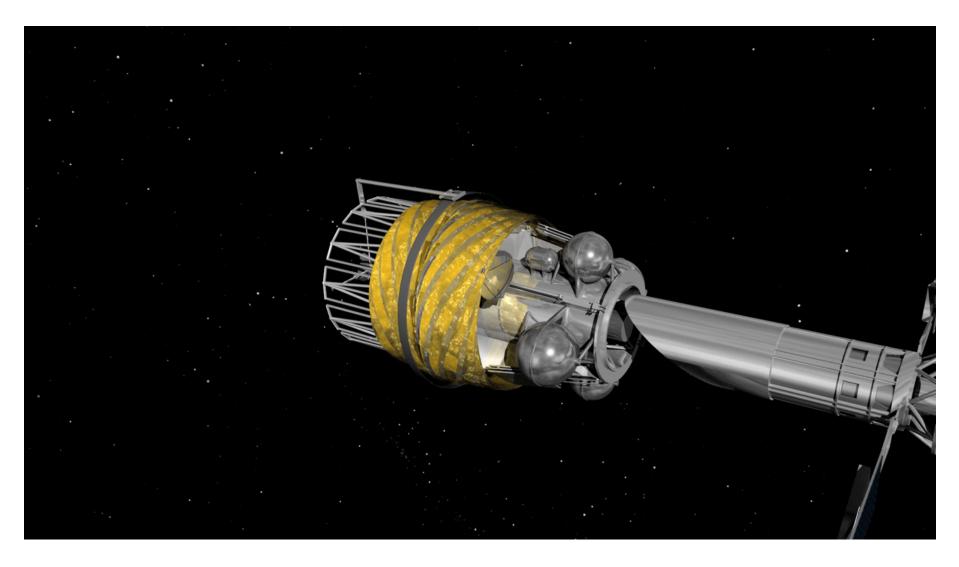
## Probe STDT Progress - Starshade



**STDT:** S. Seager (Chair, MIT), W. Cash (Colorado), S. Domagal-Goldman (NASA GSFC), N. J. Kasdin (Princeton), M. Kuchner (NASA GSFC), A. Roberge (NASA GSFC), S. Shaklan (NASA JPL), W. Sparks (STScI), M. Thomson (NASA JPL), M. Turnbull (GSI)

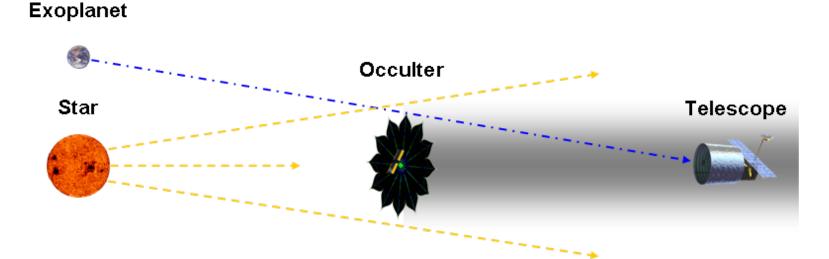
Design Team: D. Lisman, E. Cady, S. Martin, D. Webb (NASA JPL)

## Starshade Introduction



# Starshade Strengths

- Contrast and inner working angle is decoupled from the telescope aperture size
  - A "simple" space telescope can be used
  - No wavefront correction is needed
- No outer working angle



## Starshade Strengths

- High quality telescope is not required
  - Segments and obstructions are not a problem
- High throughput, broad bandpass
- Inner working angle can be changed by altering the telescope-starshade separation
- 360 degree suppressed field of view
- No constraints on other astronomical instruments

# Starshade Challenges

- Full scale end-to-end system test on the ground is not possible
  - Subscale lab and field tests are ongoing



T. Glassman / NGAS

- Limited number of starshade movements
- Long durations between observations while moving the starshade

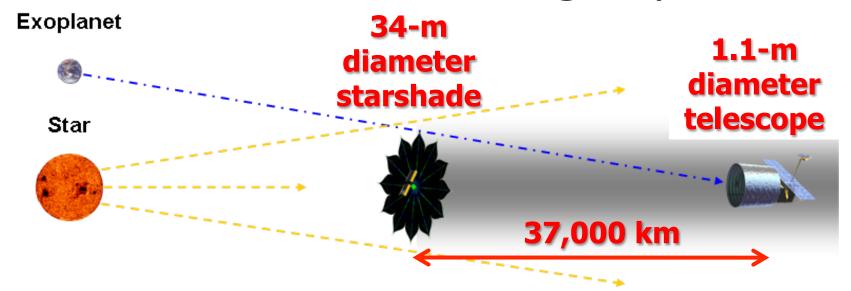
# Starshade Challenges

- On orbit deployment of a large structure
- Precise edge profile (~ 50 μm tolerance) required over large structure
- Precise alignment between starshade and telescope (i.e., formation flying) to ± 1 meter tolerance
- Costing



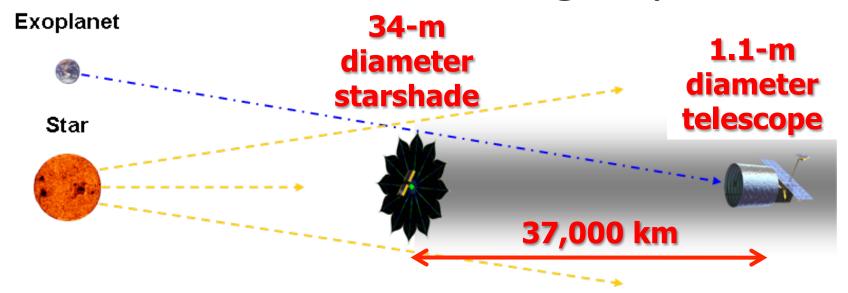
NASA / JPL / Princeton

## Probe Baseline Design Specs



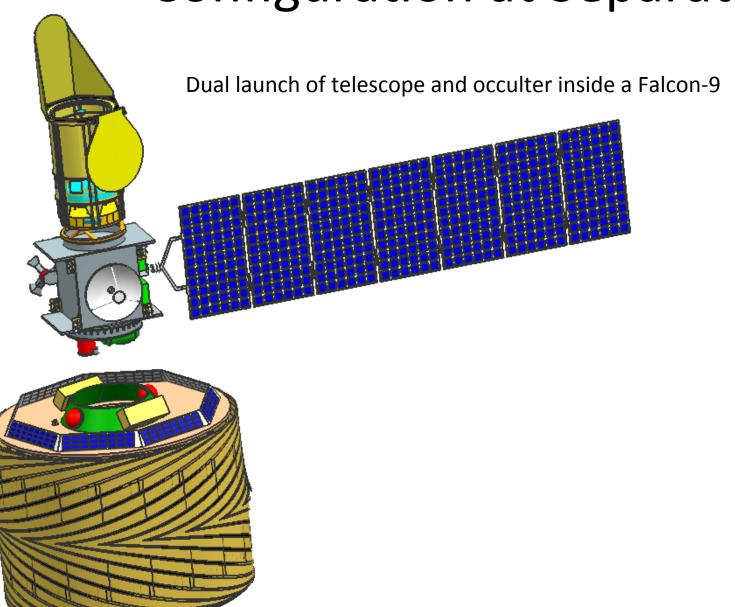
- Off-the-shelf on-axis optical telescope (1.1 m NextView)
- Earth-leading orbit
- Move telescope, not starshade for retargeting
- Instrumentation: imager and low resolution spectrograph

## Probe Baseline Design Specs



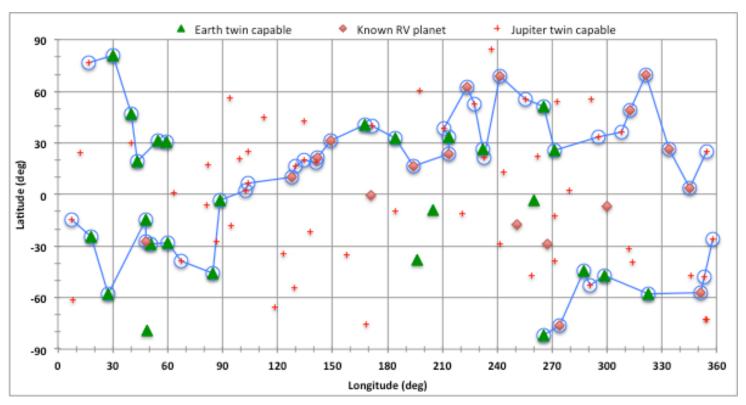
- Primary operating mode
  - 500 850 nm bandpass
  - 95 milli-arcsecond inner working angle
  - Limiting fractional planet brightness  $\sim 9 \times 10^{-11}$
- Other bands with different IWAs for follow-up

# Configuration at Separation



# **Preliminary Observing Strategy**

- First 18 months in "reconnaissance mode"
  - Multi-color imaging only to find candidates
- Second 18 months for revisits and spectroscopy



D. Lisman (NASA JPL)

## **Preliminary Science Yield Predictions**

- In 18 months, observe 55 stars
  - Assuming 1 zodi of exozodi dust ...
  - Can detect Jupiter-twins around all stars
  - 14 stars with detectable known giant planets from radial velocity surveys
  - Possibility of detecting Earth-analog exoplanets around 22 stars



**NASA** 



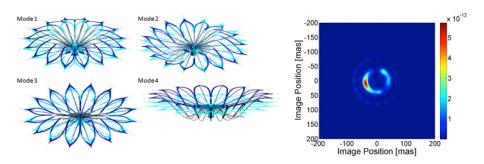
L. Cook

 Remainder of mission for revisits, follow-up spectroscopy, potentially disk observations

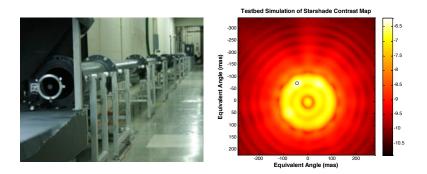
## **Technology Demonstrations**

Performance Modeling and Testing

#### Optical models with distortions



0.1% scale lab testing



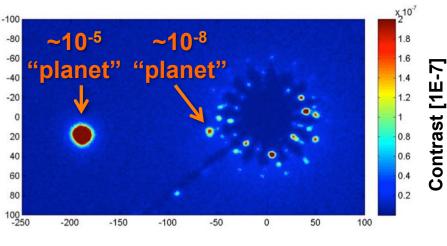
~ 1% scale field testing



T. Glassman / NGAS



Position [arcsec]

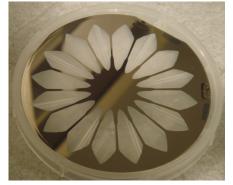


Position [arcsec]

## **Technology Demonstrations**

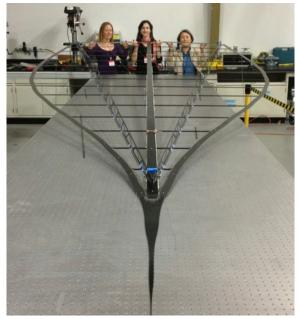
Precision petal manufacturing

#### Sub-scale full starshade



Sirbu, Kasdin, & Vanderbei 2013

Full-scale petal with required edge profile



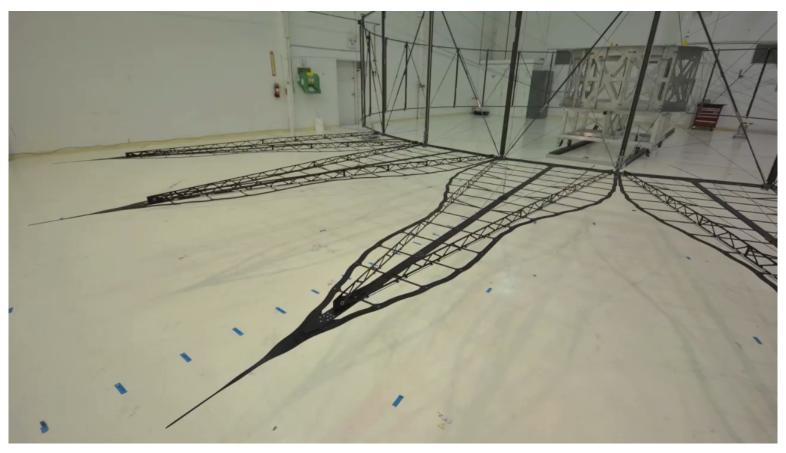


D. Lisman

Development of knife-edge to control edge scatter underway

# **Technology Demonstrations**

Starshade stowage and deployment



http://www.youtube.com/watch?v=G68sqgRhP2E

## **STDT Next Steps**

### Baseline Probe Design

- Refine Design Reference Mission and science yield simulations
- Complete trades for the baseline design of starshade + occulter system

## "Starshade Ready" Design

- Starshade design for a future or existing telescope (e.g., NRO/AFTA)
- Starshade readiness of telescope

## Technology Development

- Priorities recommended by STDT
- Technology development will continue by the community through competed
  NASA technology programs; some STDT members participating